Probability Distributions of Sonic Boom Noise Metrics in Isotropic Turbulent Fields

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- Goal: Understand how turbulence alters sonic boom metrics
- Reasons:
 - Knowledge of metrics at ground level is required to develop noise standards, plan community response surveys, etc.
 - Turbulence in the atmospheric boundary layer (ABL) randomly alters sonic boom signals
- Process: Perform numerical simulations of sonic boom through turbulence, examine probability distributions and mean values of sonic boom metrics
- Reasons:
 - Computational approaches provide a cheap alternative to flight test measurements or laboratory experiments
 - Parameters of simulated turbulent fields can be controlled

Background



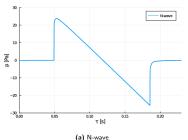
- Sonic boom metrics:
 - Stevens' Mark VII Perceived level [1] (PL)
 - Weighted Sound Exposure Levels (A-E weighted SEL)
 - ► Indoor Sonic Boom Annoyance Predictor [2] (ISBAP)
- Flight test measurements:
 - Bradley et al. [3] (SonicBAT): Average PL and ISBAP decrease with propagation distance
- Laboratory scale experiments:
 - Lipkens and Blackstock [4]: Rise time increases as N-wave propagates through turbulence
- Computational simulations:
 - Stout [5]: Average PL and ISBAP decrease in propagation direction, supported by SonicBAT measurements
- Additional background: Averiyanov et al. [6], Yuldashev et al. [7], Luquet [8]

- Simulate isotropic velocity fields on structured grid
- Numerically solve FLHOWARD equation [8] (Eqn. 1) with Strang split step [9]
- Each physical effect is computed with different numerical method (for more info see [8, 5, 10])
- ullet Extract sonic boom signals at 342 locations along the x_1 direction, and 25 locations in the transverse directions
- Compute sonic boom metrics at each microphone location for each simulation
- \bullet Compute average and standard deviations at each x_1 location across 500 simulations (12,500 microphones)

$$\frac{\partial^2 \tilde{p}}{\partial x_1 \partial \tau} = \frac{\overline{c}}{2} \frac{\partial^2 \tilde{p}}{\partial x_i \partial x_i} - \frac{u_1'}{\overline{c}} \frac{\partial^2 \tilde{p}}{\partial x_1 \partial \tau} + \frac{u_1'}{\overline{c}^2} \frac{\partial^2 \tilde{p}}{\partial \tau^2} + \frac{u_2'}{\overline{c}} \frac{\partial^2 \tilde{p}}{\partial x_2 \partial \tau} + \frac{u_3'}{\overline{c}} \frac{\partial^2 \tilde{p}}{\partial x_3 \partial \tau} + \frac{\delta}{2\overline{c}^3} \frac{\partial^3 \tilde{p}}{\partial t^3} + \frac{\beta}{2\overline{\rho}c^3} \frac{\partial^2 \tilde{p}^2}{\partial t^2}.$$
(1)

Initial Waveforms

- Two waveforms are considered:
 - N-wave
 - C609 waveform of NASA X-59 QueSST Aircraft
 - Sampled at 12.8 kHz
 - ▶ Initially propagated from aircraft to 5000 ft (7000 ft for C609) above the ground in PCBoom
- Spatial domain
 - 0 m $< x_1 \le$ 2046 m, $(\Delta x_1 = 2 \text{ m})$
 - ▶ -510 m $< x_{2,3} <$ 510 m, $(\Delta x_{2,3} = 4 \text{ m})$



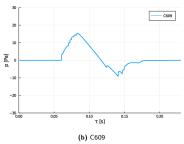


Figure 1: Initial waveforms.

Turbulent Field



- Isotropic turbulent field simulated with method of Frehlich [11]
- Temperature fluctuations are not considered
- Longitudinal integral scale L_f and rms velocity σ_u are prescribed for each propagation case considered
- N-wave and C609 signals are propagated through 500 randomly generated fields for each case

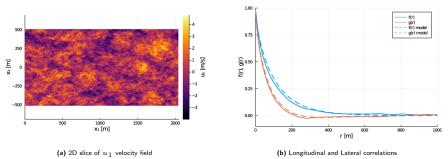


Figure 2: a) Turbulent field generated for $\sigma_u=1.0$ m/s and $L_f=100$ m. b) Comparison of average longitudinal and lateral correlations with correlation functions associated with the von Kármán spectral model.

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 According to Bradley et al. [3], turbulence intensity levels in the ABL can be categorized by the rms velocity

 $\textbf{Table 1:} \ \mathsf{Categories} \ \mathsf{of} \ \mathsf{turbulence} \ \mathsf{intensity} \ \mathsf{based} \ \mathsf{on} \ \mathsf{rms} \ \mathsf{velocity}$

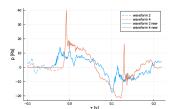
| Turbulence intensity | σ_u m/s |
|----------------------|----------------|
| Low | 0.31 |
| Medium | 1.10 |
| High | 1.89 |

- Based on these categories and other experiments [4, 6]:
 - ▶ RMS velocity: 0.2 m/s $\leq \sigma_u \leq$ 3.0 m/s
 - ▶ Integral scale: 100 m $\leq L_f \leq$ 200 m

Pre-Shock Noise



- Noise before the initial shock formed in the medium to high intensity propagation cases
- Effect of pre-shock noise was examined on a few waveforms
- Metrics of original waveforms were computed and compared to the same waveform with pre-shock noise eliminated
- Increase in PL and ISBAP. ASEL and CSEL were not altered by more than $\pm 0.6 \, dB$



| - | Waveform | Δ PL | Δ ISBAP |
|---|----------|----------------------|-----------------------|
| | 1 | 2.2×10^{-5} | -7.2×10^{-4} |
| | 2 | 0.96 | 0.72 |
| | 3 | 1.64 | 1.58 |
| | 4 | 0.09 | 0.06 |
| | 5 | 1.37 | 1.32 |

Figure 3: Two waveforms sampled from test simulations with $\sigma_u=1.8$ ${
m m/s}$ and $L_f=100$ m. The pre-shock noise that originally occured in the waveforms is shown as dashed lines.



- PL closely follows a normal distribution for lower turbulence levels
- Distribution is skewed to the right at higher turbulence levels
- Figure 4 shows distribution of PL for N-wave

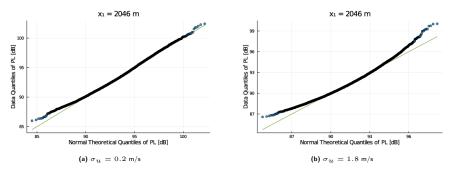


Figure 4: Quantile-Quantile plots of PLdB for the N-wave at $x_1=2046\,$ m compared to a normal distribution (solid line) for two different turbulence intensities.

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Distributions of Sonic Boom Metrics II



- N-wave:
 - Similar trend occurs for all metrics, except for ASEL and DSEL
 - ASEL is normal for low turbulence levels, skewed to the right for medium turbulence levels, and returns to normal for high turbulence levels
 - DSEL results show skewness at lower turbulence intensities
- C609 waveform:
 - All metrics follow same trend as PL, except ASEL, CSEL and DSEL
 - ASEL follows same trend as N-wave
 - CSEL remains normal for all turbulence levels
 - DSEL is normal at medium turbulence intensities, skewed to the right at low and high intensities

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Expected Values I



- Examine average PL as a function of propagation distance and turbulence intensity
- Mean PL computed from several cases conducted at low, medium, and high turbulence intensities
- Nonlinear relationship with x_1 when $x_1 \leq 500$ m
- Beyond 1000 meters, decrease in PL is approximately linear

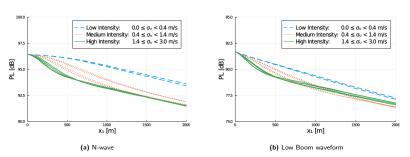


Figure 5: Average value of PL dB as a function of propagation distance for both waveforms.

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- Expected Values II
 - Model mean values beyond $x_1 = 1000$ m
 - First, examine decrease with respect to turbulence rms velocity
 - Solid lines represent exponential decay model with fitted parameters

$$PL(\sigma_u) = C_1 \left[C_3 + \exp(-C_2 \sigma_u) \right]$$
 (2)

 Exponential decay model does not capture additional dip seen in C609 waveform results and some N-wave results

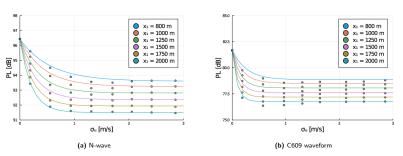


Figure 6: Average value of PLdB as a function of rms velocity for two waveforms.

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- ISBAP shows similar trend to PL
- Mismatch in the range $0.6 \le \sigma_u \le 1.8$ m/s for N-wave
- Mismatch for $0.2 \le \sigma_u \le 0.6$ m/s for C609 waveform
- Additional metrics (ASEL and CSEL) are shown in the extra slides

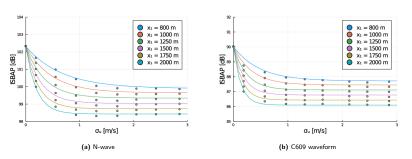


Figure 7: Average ISBAP as a function of σ_u for a) N-wave and b) C609, compared to exponential decay model.

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Summary and Conclusions



- Computational results suggest the following:
 - Sonic boom metrics are normally distributed at low turbulence intensities
 - As σ_u increases, distributions become skewed to the right
 - Average values of PL and ISBAP are approximately constant with respect to σ_u when $\sigma_u > 2$ m/s
 - ▶ The decrease in average PL and ISBAP with respect to σ_u can be approximated by an exponential decay curve
- Future work:
 - Resolve issues with pre-shock noise
 - Investigate changes to sonic boom metrics when integral scale is increased
 - Present data for standard deviations and 95% confidence intervals in future publication
 - Propagation through inhomogeneous turbulence



Thank you

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References I



- S. S. Stevens, "Perceived level of noise by mark VII and decibels (e)," The Journal of the Acoustical Society of America, vol. 51, pp. 575–601, Feb. 1972.
- J. Rathsam, A. Loubeau, and J. Klos, "A study in a new test facility on indoor annoyance caused by sonic booms," Tech. Rep. NASA/TM-2012-217332, National Aeronautics and Space Administration, Langley Research Center, 2012.
- [3] K. A. Bradley, C. M. Hobbs, C. B. Wilmer, V. W. Sparrow, T. A. Stout, J. M. Morgenstern, K. H. Underwood, D. J. Maglieri, R. A. Cowart, M. T. Collmar, and et al., "Sonic booms in atmospheric turbulence (sonicbat): The influence of turbulence on shaped sonic booms," NASA Technical Report, 2020.
- [4] B. Lipkens and D. T. Blackstock, "Model experiment to study sonic boom propagation through turbulence. part i: General results," The Journal of the Acoustical Society of America, vol. 103, pp. 148–158, Jan. 1998.
- T. Stout, Simulation of N-wave and shaped supersonic signature turbulent variations. PhD thesis. Pennsylvania State University. Dec. 2018.
- [6] M. Averiyanov, P. Blanc-Benon, R. O. Cleveland, and V. Khokhlova, "Nonlinear and diffraction effects in propagation of N-waves in randomly inhomogeneous moving media," *The Journal of the Acoustical Society of America*, vol. 129, pp. 1760–1772, Apr. 2011.
- [7] P. V. Yuldashev, S. Ollivier, M. M. Karzova, V. A. Khokhlova, and P. Blanc-Benon, "Statistics of peak overpressure and shock steepness for linear and nonlinear n-wave propagation in a kinematic turbulence," The Journal of the Acoustical Society of America, vol. 142, pp. 3402–3415, Dec. 2017.
- [8] D. Luquet, 3D simulation of acoustical shock waves propagation through a turbulent atmosphere. Application to sonic boom.
 Theses. Université Pierre et Marie Curie Paris VI. Jan. 2016.
- [9] G. Strang, "On the construction and comparison of difference schemes," SIAM Journal on Numerical Analysis, vol. 5, pp. 506–517, Sept. 1968.
- [10] A. N. Carr, J. B. Lonzaga, and S. A. Miller, "Modelling sonic boom propagation through planetary boundary layer turbulence near the lateral extent of the carpet," The Journal of the Acoustical Society of America, vol. 148, pp. 2739–2739, Oct. 2020.
- [11] R. Frehlich, L. Cornman, and R. Sharman, "Simulation of three-dimensional turbulent velocity fields," *Journal of Applied Meteorology*, vol. 40, pp. 246–258. Feb. 2001.

Cases I



• These are all of the cases conducted for both N-wave and C609

Table 2: Test Cases for Propagation through Isotropic Turbulence (Suite 1)

| Test Cases | σ m/s | L_f m | d_1 | $d_{2,3}$ | Δx_1 | $\Delta x_{2,3}$ | f_s kHz |
|------------|--------------|---------|--------|-----------|--------------|------------------|-----------|
| Case 1 | 0.0 | 0.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 2 | 0.2 | 100.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 3 | 0.6 | 100.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 4 | 1.0 | 100.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 5 | 1.4 | 100.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 6 | 1.8 | 100.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 7 | 2.2 | 100.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 8 | 2.6 | 100.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 9 | 3.0 | 100.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |

Cases II



Table 3: Test Cases for Propagation through Isotropic Turbulence (Suite 3)

| Test Cases | σ m/s | L_f m | d_1 | $d_{2,3}$ | Δx_1 | $\Delta x_{2,3}$ | f_s kHz |
|------------|--------------|---------|--------|-----------|--------------|------------------|-----------|
| Case 18 | 0.2 | 150.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 19 | 0.6 | 150.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 20 | 1.0 | 150.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 21 | 1.4 | 150.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 22 | 1.8 | 150.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 23 | 2.2 | 150.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 24 | 2.6 | 150.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 25 | 3.0 | 150.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |

Cases III



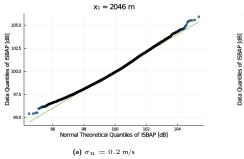
Table 4: Test Cases for Propagation through Isotropic Turbulence (Suite 5)

| Test Cases | σ m/s | L_f m | d_1 | $d_{2,3}$ | Δx_1 | $\Delta x_{2,3}$ | f_s kHz |
|------------|--------------|---------|--------|-----------|--------------|------------------|-----------|
| Case 34 | 0.2 | 200.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 35 | 0.6 | 200.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 36 | 1.0 | 200.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 37 | 1.4 | 200.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 38 | 1.8 | 200.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 39 | 2.2 | 200.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 40 | 2.6 | 200.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |
| Case 41 | 3.0 | 200.0 | 2046 m | 1020 m | 2 m | 4 m | 12.8 |

ISBAP N-wave



- Similar trend to PL
- ISBAP is skewed to the right at higher turbulence levels



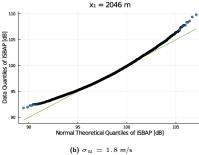


Figure 8: Quantile-Quantile plots of ISBAP dB for the N-wave at $x_1=2046\,$ m compared to a normal distribution (solid line) for two different turbulence intensities.

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ASEL N-wave



- ASEL differs from PL because it follows a normal distribution at high turbulence levels
- At $\sigma_u=3.0$ m/s ASEL is much closer to the theoretical normal distribution than other metrics

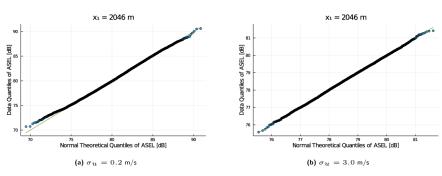


Figure 9: Quantile-Quantile plots of ASEL dB for the N-wave at $x_1=2046\,\mathrm{m}$ compared to a normal distribution (solid line) for two different turbulence intensities.

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BSEL N-wave



- Similar trend to PL
- BSEL is skewed to the right at higher turbulence levels

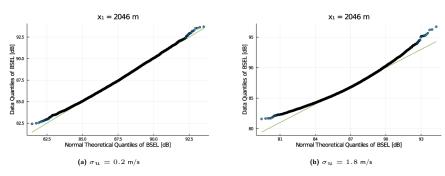


Figure 10: Quantile-Quantile plots of BSEL dB for the N-wave at $x_1\,=\,2046\,$ m compared to a normal distribution (solid line) for two different turbulence intensities.

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CSEL N-wave



- Similar trend to PL
- CSEL is skewed to the right at higher turbulence levels

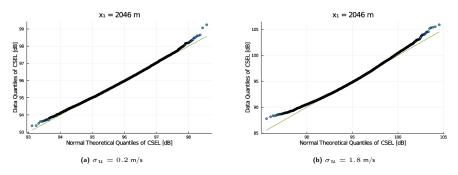


Figure 11: Quantile-Quantile plots of CSEL dB for the N-wave at $x_1 = 2046$ m compared to a normal distribution (solid line) for two different turbulence intensities.

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DSEL is skewed to the right at all turbulence levels

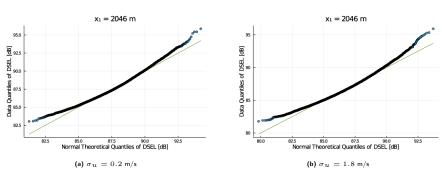
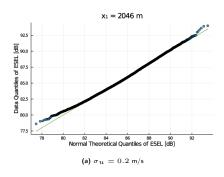


Figure 12: Quantile-Quantile plots of DSEL dB for the N-wave at $x_1\,=\,2046$ m compared to a normal distribution (solid line) for two different turbulence intensities.

ESEL N-wave



- Similar trend to PL
- ESEL is skewed to the right at higher turbulence levels



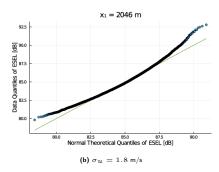


Figure 13: Quantile-Quantile plots of ESEL dB for the N-wave at $x_1=2046$ m compared to a normal distribution (solid line) for two different turbulence intensities.

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PL C609 Waveform

- Similar trend of PL for the N-wave
- PL is skewed to the right at higher turbulence levels
- Figure 14 for C609 waveform at two turbulence intensities

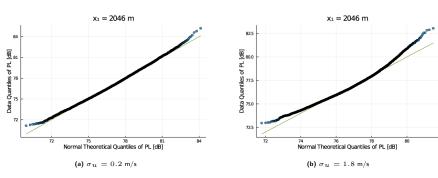


Figure 14: Quantile-Quantile plots of PL dB for the C609 waveform at $x_1=2046$ m compared to a normal distribution (solid line) for two different turbulence intensities.

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ISBAP C609 Waveform



- Similar trend to PL
- ISBAP is skewed to the right at higher turbulence levels

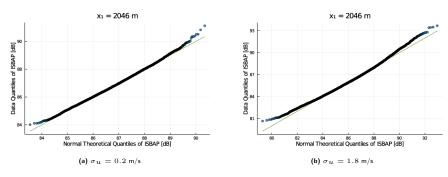


Figure 15: Quantile-Quantile plots of ISBAP dB for the C609 Waveform at $x_1=2046$ m compared to a normal distribution (solid line) for two different turbulence intensities.

ASEL C609 Waveform



- ASEL differs from PL because it follows a normal distribution at high turbulence levels
- At $\sigma_u = 1.8$ m/s ASEL only has a slight skewness, and is much closer to the theoretical normal distribution than other metrics

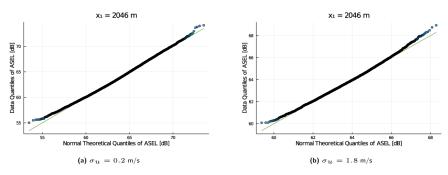


Figure 16: Quantile-Quantile plots of ASEL dB for the C609 Waveform at $x_1=2046$ m compared to a normal distribution (solid line) for two different turbulence intensities

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- Similar trend to PL
- BSEL is skewed to the right at higher turbulence levels

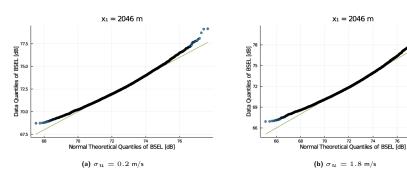


Figure 17: Quantile-Quantile plots of BSEL dB for the C609 Waveform at $x_1=2046\,\mathrm{m}$ compared to a normal distribution (solid line) for two different turbulence intensities.

CSEL C609 Waveform



- Similar trend to PL
- CSEL is skewed to the right at higher turbulence levels

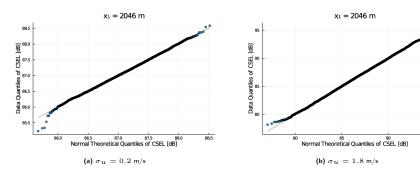


Figure 18: Quantile-Quantile plots of CSEL dB for the C609 Waveform at $x_1=2046$ m compared to a normal distribution (solid line) for two different turbulence intensities.

DSEL C609 Waveform

 DSEL is normal at medium turbulence intensities and skewed for low and high turbulence

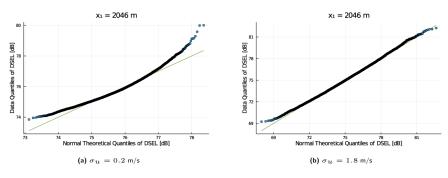


Figure 19: Quantile-Quantile plots of DSEL dB for the C609 Waveform at $x_1=2046\,\mathrm{m}$ compared to a normal distribution (solid line) for two different turbulence intensities.

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ESEL C609 Waveform



75.0

- Similar trend to PL
- ESEL is skewed to the right at higher turbulence levels

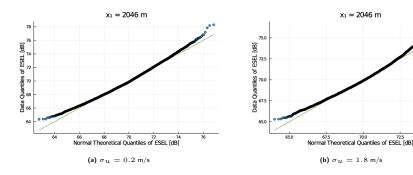


Figure 20: Quantile-Quantile plots of ESEL dB for the C609 Waveform at $x_1=2046$ m compared to a normal distribution (solid line) for two different turbulence intensities.

Average ISBAP I

- ISBAP follows same trend as PL with respect to increasing turbulence intensity
- Can be approximated with exponential decay model
- Not valid at low turbulence intensity for the C609 waveform

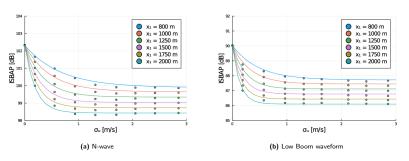


Figure 21: Average ISBAP as a function of σ_{th} for a) N-wave and b) C609, compared to exponential decay model.



- Same trends as PL
- ullet Approximately linear for $x_1>1000$ m for low and high turbulence intensities

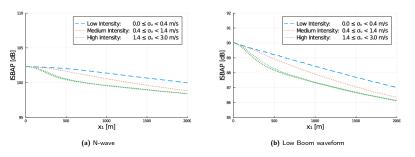


Figure 22: Average ISBAP as a function of x_1 for a) N-wave and b) C609.

Average ASEL



- ASEL follows same trend as PL for the N-wave
- For the C609 waveform, there is a significant drop in ASEL for medium turbulence levels
- At higher turbulence levels, the ASEL metric approaches a constant value

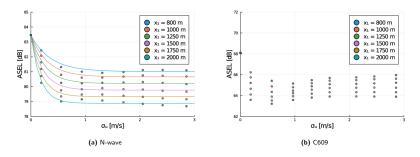


Figure 23: Average ASEL as a function of σ_{1L} for the N-wave and C609 waveform.

- CSEL does not follow PL or ISBAP trends
- For C609 waveform, CSEL is monotonically decreasing
- \bullet For the N-wave, CSEL approaches a constant value when $\sigma_u > 2.0~{\rm m/s}$

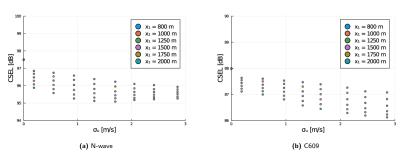


Figure 24: Average CSEL as a function of σ_u for the N-wave and C609 waveform.